



# **A 3D-CTM with detailed online PSC-microphysics: analysis of the Antarctic winter 2003 by comparison with satellite observations**

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# A 3D-CTM with detailed online PSC-microphysics: analysis of the Antarctic winter 2003 by comparison with satellite observations

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## Abstract

We present the first detailed microphysical simulations which are performed online within the framework of a global 3-D chemical transport model (CTM) with full chemistry. The model describes the formation and evolution of four types of polar stratospheric cloud (PSC) particles. Aerosol freezing and other relevant microphysical processes are treated in a full explicit way. Each particle type is described by a binned size distribution for the number density and chemical composition. This set-up allows for an accurate treatment of sedimentation and for detailed calculation of surface area densities and optical properties. Simulations are presented for the Antarctic winter of 2003 and comparisons are made to a diverse set of satellite observations (optical and chemical measurements of POAM III and MIPAS) to illustrate the capabilities of the model. This study shows that a combined resolution approach where microphysical processes are simulated in coarse-grained conditions gives good results for PSC formation and its large-scale effect on the chemical environment through processes such as denitrification, dehydration and ozone loss. It is also shown that the influence of microphysical parameters can be measured directly from these processes.

## 1 Introduction

Polar stratospheric clouds (PSCs) play a double key role in springtime polar ozone depletion, see e.g. Solomon (1999); WMO (1999); Dessler (2000); Tolbert and Toon (2001) and references therein. The surface of the PSC particles serves as a catalyzing substrate for heterogenous reactions which transfer chlorine from reservoir species to active species. Upon photolysis these active species release atomic chlorine which in its turn ignites catalytic ozone destruction cycles. The process of chlorine activation would be severely underestimated without taking into account the catalyzing role of PSCs.

The second key role of PSCs in ozone depletion is indirect in the way that they control

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the nitrogen budget of the polar stratosphere, at first temporarily through the uptake of nitric acid, a process called denoxification. Then when the particles grow large enough to reach substantial fall velocities, they will sediment and permanently remove the absorbed nitrogen from the stratosphere, this process is called denitrification. Such a situation in which the polar stratosphere becomes permanently depleted from nitrogen throughout the polar winter and spring will delay the deactivation of active chlorine back into the reservoir species and as a consequence prolong the ozone depletion conditions.

As liquid PSC particles (Supercooled Ternary Solutions, STS) are not expected to grow very large the current general view is that only solid particles can cause denitrification. Besides ice particles, which freeze out of STS at very low temperatures, nitric acid trihydrate (NAT) has progressively been regarded as a second important type of solid PSC particle (Voigt et al., 2000; Fahey et al., 2001). Not only is NAT a remnant of evaporation of ice particles, but recent research indicates that NAT can also form out of metastable nitric acid dihydrate (NAD) which may freeze out of STS at temperatures above the ice frost point (Tabazadeh et al., 2002). This means that NAT may nucleate and grow at higher temperatures than ice.

In the past various models have been developed to understand how PSCs form and evolve and to study their influence on the chemical environment. Most of these models are defined as Lagrangian box models, e.g. the DMI microphysical model used in the present study (Larsen, 2000) has been extensively used for studies of PSC evolution on trajectories (Larsen et al., 2002, 2004; Svendsen et al., 2002; Höpfner et al., 2006a). The IMPACT model (Drdla, 1996; Drdla et al., 2003) couples a detailed microphysical description and a full chemistry approach on trajectories. Various models estimate the effect of PSCs on the chemical fields by using a simplified growth and sedimentation scheme. The CLaMS model (Groß et al., 2002; Konopka et al., 2004) follows such an approach combined with a full chemical description on trajectories. A difficulty with trajectory models concerns the handling of sedimentation. In the mentioned models various approaches have been used to deal with this problem. The most common

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solution consists in determining the fall velocity of the largest particles and deriving from this a downward flux for  $\text{H}_2\text{O}$  and  $\text{HNO}_3$ , eventually combined with a parametrization for particle evaporation (Grooß et al., 2004). The IMPACT model uses a more complex approach based on the flexible grid concept of Müller and Peter (1992). A different approach aimed specifically at the study of denitrification is used in Carslaw et al. (2002). In this study a denitrification scheme (DLAPSE) is defined on trajectories in which the sedimentation of particles is treated analytically. This scheme is coupled to the SLIMCAT 3-D chemical transport model (CTM) (Chipperfield, 1999). This approach has been successful in determining the magnitude and spatial distribution of denitrification in several Arctic winters (Mann et al., 2003; Davies et al., 2005). More simplified efforts to study the effect of PSCs on chlorine and hence ozone chemistry are used by many 3D-CTMs which incorporate PSC parametrizations based on the thermodynamical conditions. Such approaches seem to be sufficient for what concerns the large-scale effect of PSCs on the chemistry. A problem is that important processes such as denitrification have to be parametrized in some way in such models.

The model presented in this paper is, to the best of our knowledge, the first Eulerian 3D-CTM to incorporate detailed PSC microphysics, including an explicit treatment of sedimentation. It is a more generalized continuation of the work of Fonteyn and Larsen (1996) which studied PSC formation in a 2-D context. The microphysical model we use studies ensembles of particles in detail by following their evolution through a fixed binned size distribution in which the number densities and the chemical composition of all particle types are stored. This detailed microscopical information allows for the explicit calculation of various quantities both microscopic (e.g. surface area densities) and macroscopic (such as optical properties of clouds, e.g. extinction). The binned ensemble of particles is treated as a full part of the CTM and is advected in a similar way as the chemical gas-phase species. The bin structure also allows for an explicit description of sedimentation. Indeed, for each size bin the fall velocity can be calculated and hence the number of particles falling out of that bin over one model time step. Because fixed size bins are used in the model, the sedimenting particles can take their

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appropriate place in the size bins at the lower vertical model level.

All this makes it possible to compare the model output to various simultaneous observational datasets as diverse as optical properties and chemical concentrations. Because the model treats ice as well as NAT particles in a general way, it is applicable to both Arctic and Antarctic winters. In this paper we will present results for the Antarctic winter of 2003, and compare them to observations of MIPAS/Envisat (tracer evolution, denitrification, dehydration and ozone depletion) and POAM III (aerosol and PSC extinction, dehydration, ozone depletion). Results for Arctic winters will be presented in future publications.

Section 2 resumes the key facts of the PSC and CTM models. In Sect. 3 the results for the Antarctic winter 2003 will be presented, followed by a summary in Sect. 4.

## 2 The PSC and CTM models

The coupled PSC-CTM model is developed at BIRA-IASB. This model is the core model of the fourdimensional variational (4D-VAR) chemical data assimilation system BASCOE, the Belgian Assimilation System of Chemical Observations from Envisat (Errera and Fonteyn, 2001; Fonteyn et al., 2002, 2004). The PSC part is described in Larsen (2000). Here we will briefly summarize the main characteristics of the model.

### 2.1 The 3D-CTM

This 3D-CTM integrates in time the volume mixing ratios of 57 gas-phase species relevant for stratospheric chemistry. The horizontal resolution is variable and for the present study two resolutions have been used:

- low resolution:  $3.75^\circ$  in latitude and  $5^\circ$  in longitude (with  $\Delta t=30$  min.)
- high resolution:  $1.875^\circ$  in latitude and  $2.5^\circ$  in longitude (with  $\Delta t=15$  min.)

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This means that for the higher resolution the latitudinal size of a model grid cell is about 200 km and the longitudinal size ranges in the polar region from 140 km at 60° to e.g. 25 km at 85°. (We will comment later in this paper on the influence of model resolution.) The model is defined on 37 vertical levels, of which the 28 upper levels are identical to the ECMWF stratospheric pressure levels and the remaining 9 levels are a subset of the ECMWF hybrid tropospheric levels (in the 60-level product). The model's vertical range is from 0.1 hPa down to the surface. The advection scheme is the flux form semi-Lagrangian transport algorithm of [Lin and Rood \(1996\)](#). The model is driven by the 12 to 36 h ECMWF operational forecasts, which are interpolated to a 1°×1° grid before downloading, and finally averaged in a mass-conservative way to the relevant CTM grid. The choice for using forecast fields was inspired by [Meijer et al. \(2004\)](#) in which forecast fields are shown to be less diffusive than 4D-VAR analyses. Since we intend to do runs of about 6 months, we opted for the forecasts. The model integration timestep is 30 min in the low resolution and 15 min in the high resolution case. The chemistry module is built by the Kinetic PreProcessor ([Damian et al., 2002](#)) and is integrated using a third-order Rosenbrock solver ([Hairer and Wanner, 1996](#)). The 57 chemical species interact through 143 gas-phase reactions, 48 photolysis reactions and 9 heterogeneous reactions, all listed in the latest Jet Propulsion Laboratory compilation ([Sander et al., 2003](#)). It is important to note that the chemistry module includes the heterogeneous reactions and that photochemical, gas-phase and heterogeneous reactions are solved as one chemical system, in which the surface area densities are provided by the microphysical module.

## 2.2 The PSC module

The PSC module of DMI ([Larsen, 2000](#)) is interactively coupled to the 3D-CTM. It describes the evolution in size distribution of the number density and chemical composition of 4 types of particles, listed in Table 1. The size distribution of all particle types is binned on a geometrically increasing volume scale. For the present study the range in particle radius of the bins is 0.002–36 μm and the number of bins is 36. In each

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bin 4 quantities are stored: the particle number density and the masses per particle of respectively condensed sulfuric acid, nitric acid and water. Besides the ambient air temperature and pressure, the PSC module is called with the model partial pressures of water vapor and nitric acid as inputs. These model fields can be changed by the PSC module, as the particles may take up or release water or nitric acid as a consequence of their microphysical evolution. The combined surface area density of all particles on a model gridpoint is used to calculate the heterogeneous reaction rates which are served as an input to the chemistry module. This illustrates the interactive nature of the coupling between the 3D-CTM and the PSC module, which is schematically presented in Fig. 1. The PSC module has an internal timestep because the microphysical processes may require computational timescales much smaller than the CTM timestep. This internal timestep is made variable to gain computing time and is determined by the smallest timescale involved in the microphysical processes relevant in a specific call to the module.

The microphysical processes described by the module are, schematically (see [Larsen \(2000\)](#) for a detailed discussion):

1. homogeneous volume dependent nucleation of ICE out of STS 3–4 K below the ice frost point temperature  $T_{ice}$  ([Koop et al., 2000](#)), and homogeneous surface dependent nucleation of nitric acid dihydrate (NAD) out of STS ([Tabazadeh et al., 2002](#)) above  $T_{ice}$ , followed by (assumed) instantaneous conversion of NAD to NAT.
2. nucleation of NAT by vapor deposition of  $\text{HNO}_3$  and  $\text{H}_2\text{O}$  on pre-activated SAT ([Zhang et al., 1996](#)), and nucleation of ICE by vapor deposition of  $\text{H}_2\text{O}$  on NAT;
3. dissolution of SAT into STS at low temperatures in the presence of high  $\text{HNO}_3$  concentrations in the gas phase, followed by uptake of  $\text{H}_2\text{O}$  and  $\text{HNO}_3$  ([Koop and Carslaw, 1996](#)), and melting of SAT into liquid sulfate aerosols particles at temperatures above 216 K;

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4. condensation and evaporation of  $\text{HNO}_3$  and  $\text{H}_2\text{O}$  to and from STS, NAT and ICE, using the basic vapor diffusion equation and applying a full kinetic approach;
5. mass balance calculations of  $\text{HNO}_3$  and  $\text{H}_2\text{O}$  between gas and condensed phase;
6. sedimentation of particles (Pruppacher and Klett, 1997; Fuchs, 1964). As the size bins are the same everywhere in the model, particles sedimenting from one model layer to the lower one can fall into the appropriate size bin.

Recent studies seem to indicate that the homogeneous freezing rates for NAD have to be reduced considerably (Larsen et al., 2004; Irie et al., 2004). We will investigate the effect of reducing the theoretical homogeneous freezing rates by 100.

To compare with extinction measurements the extinction of the aerosols and PSCs are calculated in the model. This is done using by Mie scattering theory for the spherical particles (STS) and the T-matrix technique (Mishchenko and Travis, 1998) for the nonspherical particles (SAT, NAT, ICE). For the T-matrix calculations a database of expansions of the elements of the scattering matrix in generalised spherical functions is used. The recommendations of Mishchenko and Travis (1998) for input parameter settings have been followed. The database used is for volume-equivalent sizes and an imaginary refractive index of  $10^{-8}$ . The real refractive indices used are taken from Krieger et al. (2000) and Scarchilli et al. (2005). For the nonspherical particles an aspect ratio of 1.05 is assumed.

### 3 Antarctic winter 2003

In this section we present results of simulations for the Antarctic winter of 2003 and comparisons to satellite observations.

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### 3.1 Observational datasets and methodology

For this study of the Antarctic winter 2003 we will compare model results to observations from MIPAS/Envisat and POAM III. Here we will shortly present the instruments and their datasets, and the methodology followed for the comparisons.

5 MIPAS, or the Michelson Interferometer for Passive Atmospheric Sounding, is a limb-scanning Fourier transform infrared spectrophotometer (Fischer and Oelhaf, 1996; ESA, 2000) onboard Envisat. Envisat's orbit time is 101 min resulting in about 14 orbits per day. There are usually around 1000 MIPAS profiles available per observed species per day. Each profile has at most 17 points ranging from about 6 km to about 60 km.  
10 The chemical species observed by MIPAS are  $O_3$ ,  $HNO_3$ ,  $H_2O$ ,  $NO_2$ ,  $N_2O$  and  $CH_4$ . We use the offline reprocessed data version 4.61 for our comparisons.

POAM III is a visible/near infrared solar occultation photometer onboard the French Spot-4 satellite which measures the chemical stratospheric constituents  $O_3$ ,  $H_2O$  and  $NO_2$ , and aerosol extinction in the polar regions (Lucke et al., 1999; Lumpe et al.,  
15 2002). Here we use the POAM III version 4 data. Throughout this text we will denote POAM III by POAM. POAM profiles have a vertical resolution of 1 km. The aerosol extinction profiles reach up to 25 km. The chemical species profiles have a wider altitude range. The POAM line of sight is about 230 km. POAM measures at most 15 profiles per day in each hemisphere, all these profiles are located at a fixed latitude and with a longitude spacing of about  $25^\circ$ . The latitude at which the profiles are taken  
20 varies slowly throughout the year, but remains in the polar region. In Antarctic winter periods, the latitude varies smoothly between about  $65^\circ S$  (winter solstice) and about  $87^\circ S$  (equinox).

To provide a uniform set-up for the comparison of model results to measured profiles  
25 the following methodology will be applied. The model data have been interpolated in space and time (co-located) to the observed profiles. This interpolation occurs online, i.e. at the model timestep where a profile is available. This way of working guarantees an optimal comparison, e.g. in the case of rapidly changing PSC fields. In a practical

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sense it allows for a considerable limitation in model output and hence disk space usage. Then the observed and co-located model profiles, as well as the observational errors, are interpolated to 4 isentropic levels: 425 K, 475 K, 525 K and 575 K (corresponding to pressure values of approximately 70 hPa, 50 hPa, 30 hPa and 20 hPa and to altitudes of approximately 17.5 km, 20 km, 22.5 km and 25 km inside the vortex). Of the interpolated data at these levels the error-weighted average is calculated over 5-day intervals. Since we concentrate on polar winter processes only profiles within the polar vortex are taken into account, with the vortex edge being calculated following the definition of [Nash et al. \(1996\)](#) using the dynamical fields interpolated to the high-resolution model grid (see Sect. 2.1). Although we will call the quantities obtained in this way error-weighted mean vortex values, it is important to note that they can not be regarded as properly vortex-averaged values, but merely as averages of available values within the vortex.

When POAM samples a water ice cloud, the large solar extinction prevents POAM from performing a measurement. Such points have been removed in the corresponding model data as well before taking the average. Also observations with a negative error have been removed in both the observed and model profiles before taking the average. Negative data, e.g. some POAM extinction measurements, have not been removed because they are meaningful observations as long as their absolute value is smaller than the retrieval random error.

We have chosen to present 5-day averages because they result in much clearer plots than e.g. daily averages. For example, variability in the POAM extinction data is a convolution of the zonal asymmetry of PSC occurrence and composition, seasonal evolution of PSC forcing, and POAM sampling. Here we list a number of issues which could lead to a difference in variability between the model and the observations. POAM has a long line of sight in which various PSC particles could be sampled at various distances (within the same atmospheric layer) rather than that they would all be located at the tangent point, which is where the model calculations are done. Furthermore the model produces clouds which are limited by the model's grid size in coarse-grained

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conditions and which have uniform properties inside a grid cell. But as an occultation instrument, and even though its retrieval region is comparable to the model's horizontal grid size, POAM could well sample small localized clouds of high extinction. These will then spureously be considered as being uniformly distributed over the effective absorption region near the tangent point, leading to an increased extinction value which will be underestimated by the coarse-grained model. Also as an occultation instrument, POAM has a much higher vertical resolution than the model. And finally it is obvious that the coarse-graining of the temperature to the model resolution will have some influence. The aim of the present study is however not to analyze and eventually correct for this detailed variability (e.g. by introducing sub-grid scale mountain wave parametrizations) but rather to test the overall performance in polar winter processes of the PSC model in the coarse grained model conditions. For this 5-day averages are sufficient.

In the following we will show comparisons of model data for all observed fields except for NO<sub>2</sub>, which is nearly completely transformed into HNO<sub>3</sub> under polar winter conditions (Kawa et al., 1992), and CH<sub>4</sub>, which has a behavior comparable to N<sub>2</sub>O.

### 3.2 Initializations

The simulations start on 1 May 2003 and run until 15 December 2003. The model's chemical fields are initialized from an assimilation of MIPAS/Envisat observations with the 4D-VAR chemical data assimilation system BASCOE (Errera and Fonteyn, 2001; Fonteyn et al., 2002, 2004). An exception has been made for polar water vapor. It has been observed that there is a substantial systematic difference in polar water vapor as observed by MIPAS and by POAM, with MIPAS water lower by up to about 25% as compared to POAM in the lower stratosphere. When the model is initialized by the MIPAS water vapor in the polar regions, it underestimates the PSC extinction as measured by POAM (not shown). This is much less the case when the model's initial water vapor field is scaled in the polar regions to match the POAM measurements. Therefore we have chosen for the latter option. MIPAS v4.61 H<sub>2</sub>O data are validated so far only outside of the polarhoods (Oelhaf et al., 2004). In the recent validation

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study by Lumpe et al. (2006) it is found that POAM sunset measurements of water vapor show a systematic high bias of about 25% in the lower Antarctic stratosphere relative to HALOE, while the bias is considerably lower for the sunrise measurements. HALOE is suggested to have a 5% low bias (Harries et al., 1996; Kley et al., 2000). The latter study also shows that HALOE can show low biases of up to 15–20% with in-situ measurements. Because the differences between POAM and MIPAS are comparable to the POAM-HALOE bias, and given the suggested biases in the HALOE data, we assume that tuning the model's initial state to POAM water vapor is justifiable for the purposes of this paper. The scaling of the initial (assimilated MIPAS) water vapor field is done by deriving scaling factors from a 10-day mean POAM water vapor profile in early May 2003 and the corresponding mean model profile. Then these scaling factors are applied to the initial model state everywhere poleward of 60° latitude.

The initial sulfate aerosol size distributions are described by a lognormal function, defined by a median radius  $r_m$ , geometrical standard deviation  $\sigma$  and number density  $N_t$  (Pinnick et al., 1976). We estimated these parameters on the condition that the initial lognormal distribution should reproduce the Unified POAM reference background 1  $\mu\text{m}$  extinction monthly means (Fromm et al., 2003). This condition leaves a lot of arbitrariness in fixing the lognormal parameters. We have chosen to take a set of parameters that follow a smooth profile and deviate not too far from the measurements made by the University of Wyoming's optical particle counter (Deshler et al., 2003) on balloon flights in the Antarctic over the past decade (T. Deshler, private communication). Our estimated parameters are shown in Fig. 2. We will comment later in this paper on the influence of the choice of these parameters.

In the following we will study model runs with four different set-ups which will be denoted by:

- LR : low resolution
- LR100 : low resolution with NAD freezing rates reduced by a factor 100 (see Sect. 2.2)

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- HR : high resolution
- HR100 : high resolution with NAD freezing rates reduced by a factor 100

In the HR simulation the background aerosol distribution is updated each month except for gridpoints with temperatures below 220 K, where the microphysics is likely to play a role. In the other simulations this methodology has not been followed, and the evolution of the initial background aerosol distribution is followed throughout the winter and springtime.

### 3.3 Chemical tracers

In order to check how well the model calculates the atmospheric transport, and more specifically the adiabatic descent of air in the vortex and the dynamical isolation of the vortex, we have compared the evolution of inactive tracer species in the vortex with observational data. A badly performing transport would have significant influence on the chemical evolution. Here we compare to the temporal evolution of  $N_2O$  as observed by MIPAS following the methodology described in Sect. 3.1. The results are shown in Fig. 3.

The simulations indicate that the numerical diffusion between the  $N_2O$  depleted vortex and the surrounding areas is very large and disturbs the isolation of the vortex considerably. This numerical effect is resolution-dependent and is present in all chemical and particle fields with a large gradient over the vortex edge (e.g.  $O_3$ , PSCs). When the resolution is too low model grid cells located over the vortex edge sample regions of air both inside and outside the vortex and replace them by a single average value. This increases the values in the region where they are the lowest and vice versa (e.g. for  $N_2O$ : increase inside the vortex and decrease outside). The effect decreases with increasing resolution. In our high resolution simulations the vortex mean  $N_2O$  remains acceptable close to the observations – i.e. well within the MIPAS variability – until the second half of September. By this time most of the polar winter processes are over and the vortex starts to weaken. After this period the model basically overestimates

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the mixing of polar air when the vortex disappears. It is expected that this could be solved by using a still higher resolution but such a study was beyond the limits of the present computing power.

Results for CH<sub>4</sub> which is also measured by MIPAS are similar but the effect of numerical diffusion is smaller (not shown).

### 3.4 Extinction

Figure 4 shows the comparison of the calculated extinction at 1020 nm to the extinction as measured by POAM following the methodology described in Sect. 3.1. The agreement between model and observations is good. The correspondence is excellent on the lowest two levels. The model follows the general trend of the observations with the onset of enhanced extinction in June, the maximum in July and August, and the end of enhanced extinction in September 2003. A general discussion of the differences in variability between POAM and the model was already held in Sect. 3.1. As already mentioned there 5-day averages remove much of the small time-scale and inter-profile variability, and here they indicate that the model produces an acceptable average extinction.

On the higher levels the model predicts the onset of enhanced extinction by mid-May which is two to three weeks too early compared with the observations. It is unclear whether this problem in the early winter is due to the biases in the ECMWF temperatures or reflects a problem in the model. Gobiet et al. (2005) and Höpfner et al. (2006b) have reported considerable temperature biases in ECMWF analyses in this period (up to 3.5 K). The forecast fields which we use are initialized by these analyses and it is likely that these temperature biases propagate in the forecast fields. Reducing the NAD homogeneous freezing rates by 100 clearly improves the situation in early winter, but on the other hand leads to an underestimation of the extinction in August. When comparing 5-day median values the deviations in early winter between model and observations are even more pronounced (not shown).

The difference between HR and the other simulations in springtime is due to the

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mentioned monthly reloading of the background aerosol distribution (Sect. 3.2). The model runs with no reloading do not reproduce the actual background aerosol distribution, because there is no mechanism in the model to reload aerosols to the upper stratosphere to replace the aerosols which have been removed by PSC formation and sedimentation. This shortcoming is only playing a role at the 575 K level.

In Fig. 5a we plotted the contribution of the various particle types to the total model extinction of simulation HR100. This plot illustrates that in the model the main NAT formation period occurs in the early winter (throughout June) while STS is the major PSC type after that period, interspersed with some intermittent ICE activity, which only rarely exceeds the STS extinction. However we remind that POAM is prevented from making measurements of optically thick water ice clouds, and that the corresponding points in the co-located model profiles have been removed in the comparison, so Fig. 5a does not give an indication of the actual ICE presence. Also shown is the model's background aerosol extinction, which is initialized in May from the Unified POAM reference background. This clearly illustrates the enhancement of extinction during the winter-time due to PSC formation.

To analyze the early-winter situation a little further we have zoomed in on this period and plotted only the contributions of STS and NAT to the total model extinction in Fig. 5b for both high resolution simulations. Clearly, reducing the freezing rates decreases the NAT contribution and increases the STS contribution, as fewer NAD particles are nucleated from STS per timestep. Except at the lowest level the NAT extinction exceeds the STS contribution everywhere from mid-May onwards in the simulations with unreduced freezing rates. When the homogeneous NAD freezing rates are reduced by a factor 100 the temporal evolution of the NAT extinction corresponds very well with the evolution of the total POAM extinction. Indeed a strong enhancement in NAT extinction after the first week of June on the highest levels occurs almost simultaneously with the sudden enhancement in the POAM extinction. In the weeks before before this enhancement the overestimated extinction in the case with reduced freezing rates is caused by STS only. The STS growth and hence its extinction are controlled by

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temperature and the gasphase  $\text{H}_2\text{O}$  and  $\text{HNO}_3$  concentrations. It has been verified that initializing the simulations with the MIPAS assimilated  $\text{H}_2\text{O}$  rather than the POAM-tuned  $\text{H}_2\text{O}$  does not solve the problem completely (not shown), giving indication that temperature biases may be responsible for the problem. Simulations with an overall temperature increment of 2 K improve the situation considerably (not shown).

Now we will study the influence of the calculated PSCs on the chemical fields in the vortex.

### 3.5 Denitrification

Figure 6 shows the comparison of  $\text{HNO}_3$  between the model and MIPAS. The model overestimates the removal of nitric acid everywhere, but this overestimation is much larger on the highest two levels, while on the lowest two levels the correspondence between model and the observations is acceptable (i.e. they remain within each other's variability). The main difference between the simulations and the observations concerns the onset and rate of  $\text{HNO}_3$  removal. As shown in Tabazadeh et al. (2000) the early winter is the period when extensive denitrification occurs. The overestimation of  $\text{HNO}_3$  removal in the early winter period is likely to be a consequence of the too early onset and overestimation of enhanced PSC presence in this period, see Fig. 4. Unfortunately there are some data gaps in June exactly during the start of the denitrification process, but from the data which are available we notice an influence of the freezing rate reduction. The reduced freezing rates lead to a small delay in the onset of denitrification and seem to be closer to the observed values (although the data gap at the end of May makes it difficult to make precise conclusions). This would be consistent with the small delay in onset of enhanced PSC extinction of Fig. 4 in the same period. As a consequence the freezing rates reduction results in  $\text{HNO}_3$  values which are slightly closer to the observations during mid-June, when there are again some data available. Nevertheless the rate of denitrification in mid-June is overestimated in all simulations, which is consistent with the overestimation of the extinction in this period. By the second half of July the rate of  $\text{HNO}_3$  removal in model and observations be-

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comes comparable, but because the model has removed too much  $\text{HNO}_3$  before this date, its  $\text{HNO}_3$  values are much lower than the observed levels.

As already mentioned in Sect. 3.4 the possible influences of cold temperature biases could be responsible for this problem. The removal of  $\text{HNO}_3$  is a combination of denoxification and denitrification. Denoxification is quite sensitive to temperature biases, and an overestimated denoxification will lead to an overestimated denitrification. In the following we will try to study the process of denitrification both qualitatively and quantitatively.

As shown in e.g. Popp et al. (2001) there is a narrow correlation between the chemical tracer species  $\text{N}_2\text{O}$  and the total amount of reactive nitrogen  $\text{NO}_y$  before the start of the polar winter. In a denoxified and denitrified situation the total  $\text{NO}_y$  will deviate from this correlation with  $\text{N}_2\text{O}$ . This makes this correlation a useful tool to check the influence of PSCs on the nitrogen budget of the stratosphere. The only  $\text{NO}_y$  species measured in the standard MIPAS product are  $\text{HNO}_3$  and  $\text{NO}_2$ , but the latter is rapidly converted into  $\text{HNO}_3$  during the polar night. In fact as was demonstrated in Kawa et al. (1992) in polar winter conditions  $\text{NO}_y$  consists almost entirely of  $\text{HNO}_3$ , so for the observations of MIPAS we can approximate the  $\text{NO}_y$ – $\text{N}_2\text{O}$  correlation by a correlation between  $\text{HNO}_3$  and  $\text{N}_2\text{O}$ . In Fig. 7 we show scatterplots of  $\text{HNO}_3$  versus  $\text{N}_2\text{O}$  of MIPAS and co-located HR100 model data for four different days: 2 May, 1 July, 1 September and 1 November. The plot of 2 May illustrates that the relation of Popp et al. (2001), which was obtained for the Arctic winter of 1999–2000, is sufficiently valid for the 2003 Antarctic winter. Only above  $\sim 550\text{ K}$  the deviations between  $\text{HNO}_3$  and  $\text{NO}_y$  are visible in the model data. The effect of ongoing denitrification is clear in the plots for 1 July, and a state of nearly total denitrification is visible in the plots for 1 September. By 25 November 1 mixing of extra-vortex air with a higher potential temperature is occurring, and also the repartitioning of nitrogen within the  $\text{NO}_y$  family is illustrated in the figure.

The amount of denitrification can be calculated as the difference between  $\text{NO}_y$  under polar winter and early springtime conditions and the estimated pre-winter  $\text{NO}_y$  concentration given by the correlation with  $\text{N}_2\text{O}$ . An approximative value for denitrification as

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observed by MIPAS can then be calculated as the difference between the observed  $\text{HNO}_3$  and  $\text{NO}_y$  as derived from the observed  $\text{N}_2\text{O}$  (Davies et al., 2005). We use this approximative denitrification quantity to compare the model to the MIPAS data, see Fig. 8. The modeled denitrification remains very close to the MIPAS result on the low-est levels. On the highest levels the onset of the denitrification in the model starts too early and the initial rate of denitrification is overestimated, leading to a propagated overestimation throughout the winter. As in the comparison of  $\text{HNO}_3$  (Fig. 6) the simulations using reduced NAD freezing rates slightly delay the denitrification onset, leading to denitrification values which are closer to the observed ones in mid-June.

The deviations during late-winter and early springtime originate mainly from the problem with  $\text{N}_2\text{O}$  (see Fig. 3) due to numerical diffusion.

### 3.6 Dehydration

Figures 9 and 10 show the evolution of water vapor as observed by respectively MIPAS and POAM and the corresponding model results. The differences in polar water vapor data from both instruments which were already mentioned in Sect. 3.2 are obvious from these plots.

Dehydration occurs because water-rich ice particles sediment when they have grown sufficiently large. The MIPAS data (Fig. 9) show very little variation in the polar vortex water field throughout the winter. The POAM data of Fig. 10 show much more variability and the effect of dehydration is clearly visible. Dehydration starts slowly in June 2003, some weeks later than denitrification which is consistent with earlier studies, e.g. Tabazadeh et al. (2000). Then it continues increasingly rapidly throughout July until the beginning of August, when the polar stratospheric water vapor reaches a minimum. At this point about 60% of the initial water vapor amount has been removed. From then on the water vapor starts to increase again very slowly with the rise of temperatures, the evaporation of PSCs, the weakening of the vortex and the mixing with extra-vortex air. Substantial dehydration is only present below  $\sim 525$  K. All these results are consistent with a recent study of dehydration during the 1998 Antarctic winter using

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the IMPACT trajectory model (Benson et al., 2006).

At the 525 K level the model predicts some dehydration but this is not consistent with the observations. This level is the location of a strong vertical gradient in the water vapor field. The deviations between the model and the POAM observations may well be a consequence of the coarse vertical model resolution, which may lead to an overestimated smoothing of the water vapor gradient at this level.

Finally it is obvious that the deviations in water vapor, especially during mid-winter, are partly responsible for the underestimation in the model extinction during this period (Fig. 4).

### 3.7 Ozone depletion

Finally we have compared the evolution of model polar ozone to observations of MIPAS and POAM. The results are shown in Figs. 11 and 12. The model simulations for all set-ups reproduce the observations very well during the winter. When in the springtime the polar vortex becomes depleted with ozone due to chemical destruction, and the surrounding area, in full sunlight and in absence of active chlorine, becomes enhanced with ozone, the numerical diffusion will play an important role. During the maximum of the ozone hole conditions (second half of September and throughout October) the low resolution simulations stay about 0.5–1 ppmv above the observations. The high resolution simulations improve these deviations by ~50% at the lowest levels, while they nearly match the MIPAS data at the higher levels.

It is obvious from the plots that horizontal resolution is the main key to accurately simulating the ozone depletion of the vortex. This indicates that the PSC modelling and the related chemical processes are performing well. Good agreement is most difficult to reach at the lowest levels because the observed ozone concentrations on those levels are nearly zero, making the simulations very sensitive to numerical diffusion. It concerns here not only a problem due to the numerical diffusion of ozone alone across the vortex edge but an accumulated effect of the numerical diffusion in various fields with strong cross-vortex edge gradients such as  $\text{HNO}_3$  (Fig. 6, and the resulting

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denitrification, Fig. 8), the PSCs, and active chlorine.

Our conclusion is consistent with the study of Hoppel et al. (2005), which demonstrated that accurate dynamics is a key ingredient for correctly modeling the spatial distribution of Antarctic ozone loss.

### 5 3.8 Model sensitivities

Figures 11 and 12 illustrate that the effect on ozone depletion of numerical mixing due to horizontal resolution limitations is much larger than the microphysical effect of different NAD homogeneous freezing rates which have been studied here. The main influence of the NAD homogeneous freezing rates seems to be situated in the early winter and more specifically in the onset and rate of denitrification. Higher freezing rates imply that more NAD (and NAT) particles are formed in the same time interval so the available nitric acid will be taken up faster than in the case with lower freezing rates, where less particles will be nucleated. When in early winter the temperatures become low enough for NAD nucleation, the simulations with higher NAD freezing rates will generate more NAT particles and the denitrification will start sooner and take place faster in those simulations. When the state of nearly total denitrification is reached in July the results seem not to depend much on the NAD freezing rates anymore.

The influence of the choice of Fig. 2 of parameters for the lognormal background aerosol distribution has also been studied to some extent. The main conclusion of this study is that when the initial total number density of aerosols is too small, these few particles grow too large in size since they have all nitric acid at their disposal. The sedimentation occurs too fast and the polar stratosphere becomes nearly depleted with aerosols by mid-winter. Then from mid-winter onwards the PSC extinction is severely underestimated and as a consequence the model fails to reproduce the springtime ozone depletion.

Finally we also studied the influence of the number of size bins used to describe the the aerosol and PSC distributions. Also here we only mention the main conclusions of this test. We have increased the number of bins up to 96, but no significant improve-

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ments have been found, indicating that 36 size bins is a sufficient resolution for the studies presented here. Because even 36 size bins is computationally quite demanding (adding 4 types  $\times$  36 bins  $\times$  4 quantities = 576 extra tracers to the model) we also have checked the influence of lowering the number of size bins. We have done the simulations using 12 bins, but then the accuracy of the results decreased. For all fields still a comparable behaviour was found as in the 36 bin case, but most results were worse compared to the 36 bin case by some percents.

From the CTM point of view, a major influence on the PSC formation is the water vapor in the model. The issue of initial water vapor has been discussed before (Sect. 3.2). Initializing the model with the MIPAS water vapor which is about 20% or more lower reduces the calculated extinction by a comparable amount (not shown). As illustrated by Fig. 10, in the second half of the winter the model seems to suffer from an underestimated water vapor loading from the higher levels downwards, leading to considerable underestimated extinction values in this period, see Fig. 4. It has also been mentioned before that the biases in the ECMWF temperatures during this Antarctic winter are suspected to be a main cause of deviations found between the model simulations and the observations. In this respect we can mention the study of Benson et al. (2006) again which showed that temperature biases are the main sensitivity parameter in modelling Antarctic extinction and dehydration. We performed sensitivity tests in which the ECMWF temperatures were increased globally by 1 K and 2 K and these reduced the overestimation of the extinction in the early winter weeks considerably but still not sufficiently, which could indicate that more detailed temperature bias corrections are needed.

## 4 Summary and conclusions

We have shown that the combined approach of a detailed microphysical model running online in the coarse-grained conditions of a global CTM with full chemistry gives excellent results for polar winter processes. By simultaneously comparing PSC extinction,

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denitrification, dehydration and ozone depletion to a diverse set of observational data we have shown that this approach leads to acceptable and consistent results, leaving mainly the horizontal model resolution and temperature accuracy as key ingredients for further improvements in ozone depletion.

5 As explained before the resolution of  $1.875^{\circ} \times 2.5^{\circ}$  is our current computational limit because of the demanding microphysical module with 36 size bins. Although in the past detailed microphysics in a 3D-CTM has not been considered attractive because of the considerable computational demands, we have illustrated that such an approach is feasible. On a recent 32 CPU machine the low resolution runs take about 10' walltime  
10 per simulated day, while the high resolution runs take about 50' walltime per simulated day, which means simulating a month in the high resolution case takes about one day  
walltime.

One of the advantages of this approach is that the detailed microphysical information of the model allows for the calculation of optical properties and that in this way  
15 model results can be compared simultaneously to optical and chemical observational data. Another interesting feature of the approach is that the influence of microphysical parameters can be tested directly from the large-scale polar processes in a consistent way, as was illustrated by the influence of the homogeneous freezing rates on the denitrification.

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**Table 1.** Particle types described in the model.

Label	Description	State
STS	Sulfate aerosols, which become supercooled ternary solutions at low temperatures	Liquid
SAT	Sulfuric acid tetrahydrate particles	Solid
NAT	Nitric acid trihydrate particles	Solid
ICE	Solid ice particles	Solid

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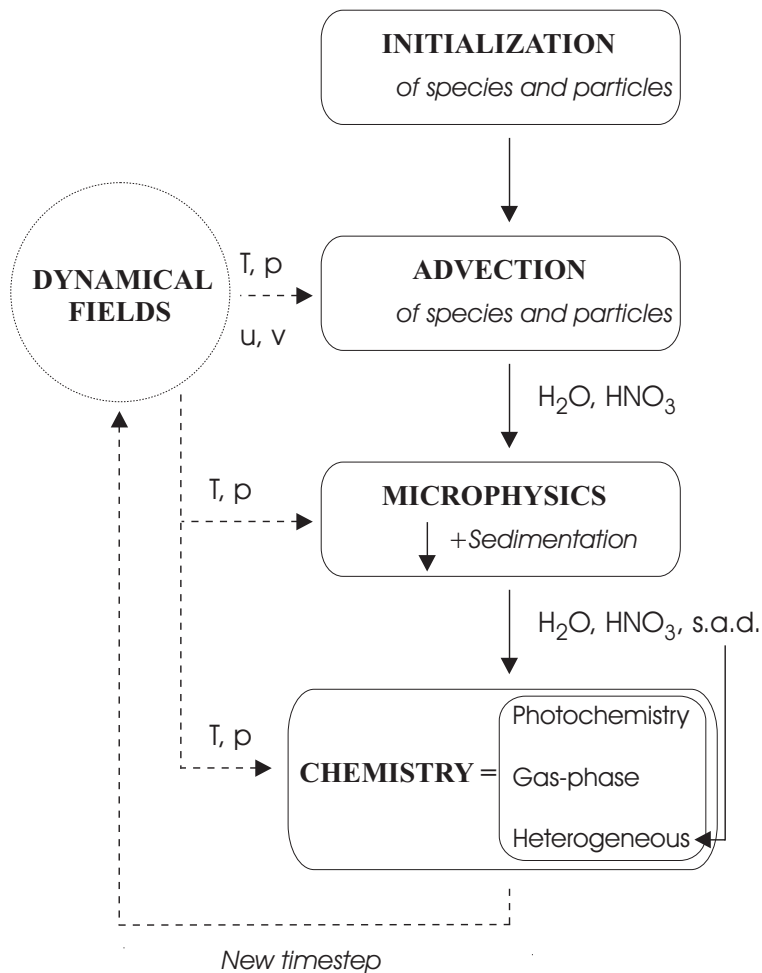
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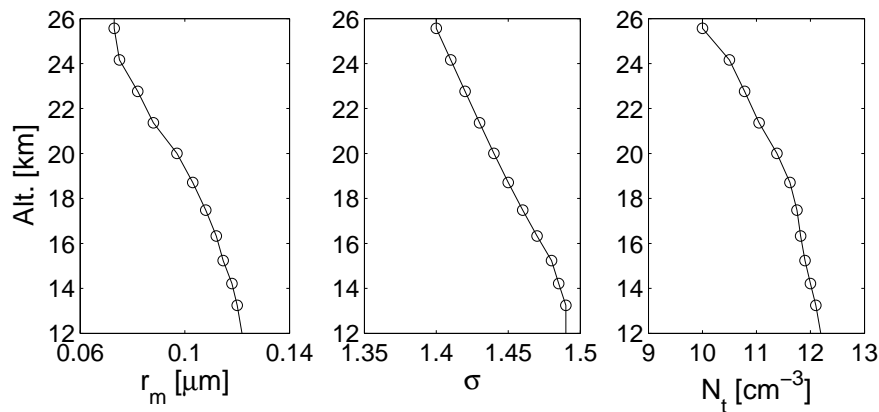


**Fig. 1.** Schematical presentation of the model, illustrating the coupling between the chemical and microphysical modules.



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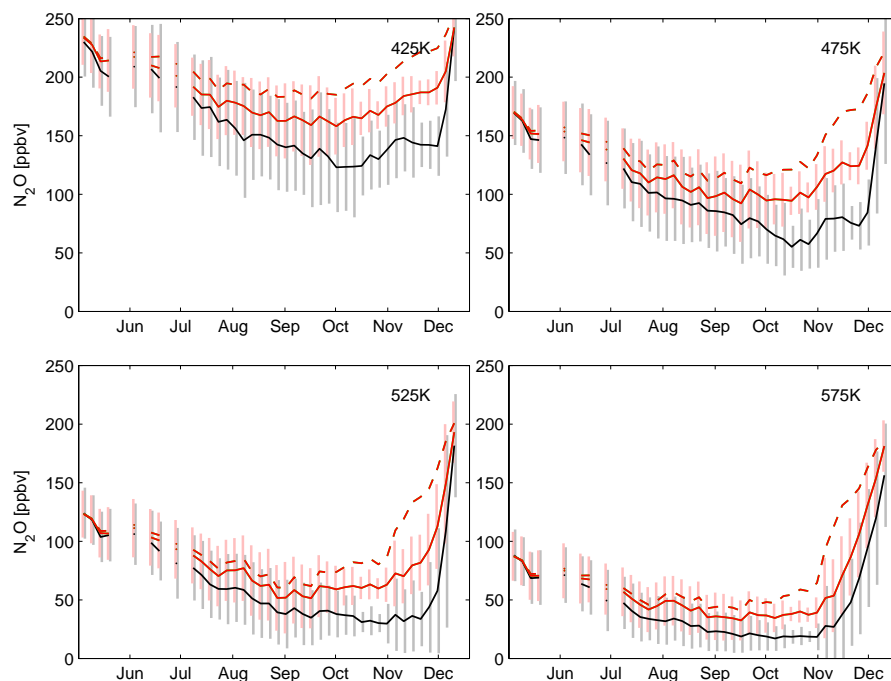


**Fig. 2.** Estimated lognormal parameters of the background aerosol size distribution in the Antarctic in May 2003 used to initialize the Antarctic simulations. From left to right: median radius  $r_m$ , geometric standard deviation  $\sigma$  and number density  $N_t$ .

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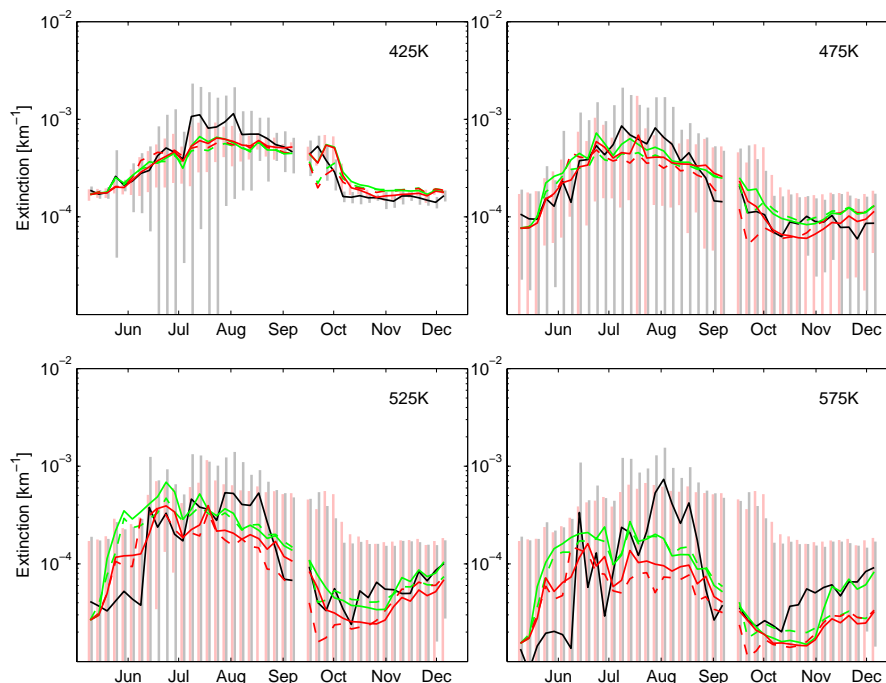


**Fig. 3.** 5-day error-weighted mean vortex  $N_2O$  mixing ratios of MIPAS (black) and co-located model results after interpolation of profiles to isentropic levels. Shown are model results for LR100 (dashed line) and HR100 (full line), they coincide with the results of respectively the LR and HR simulations. The grey-shaded area is the 5-day MIPAS variability (1 error-weighted standard deviation) while the red-shaded area is the 5-day variability of the model run HR100. For all time-series plots in this paper the horizontal axis represents months of 2003.

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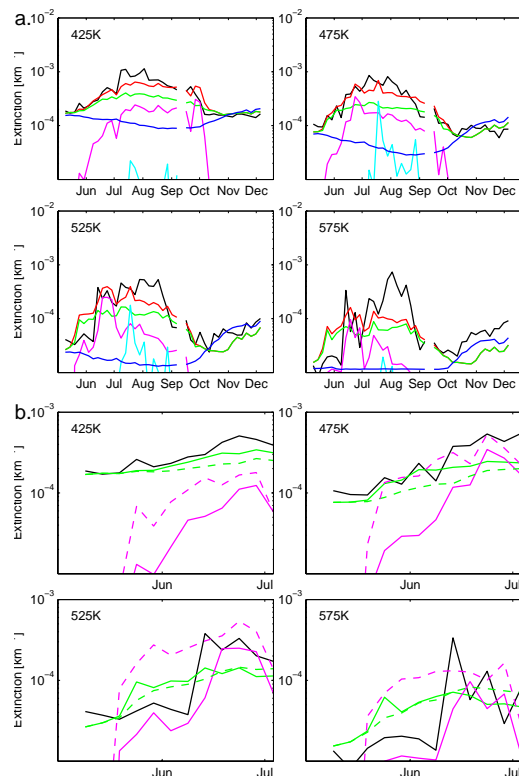


**Fig. 4.** 5-day error-weighted mean vortex 1020 nm extinction of POAM (black) and co-located model results after interpolation of profiles to isentropic levels. Model curves are: LR (green dashed), LR100 (red dashed), HR (green full), HR100 (red full). The grey-shaded area is the 5-day POAM variability (1 error-weighted standard deviation) while the red-shaded area is the 5-day variability of the model run HR100.

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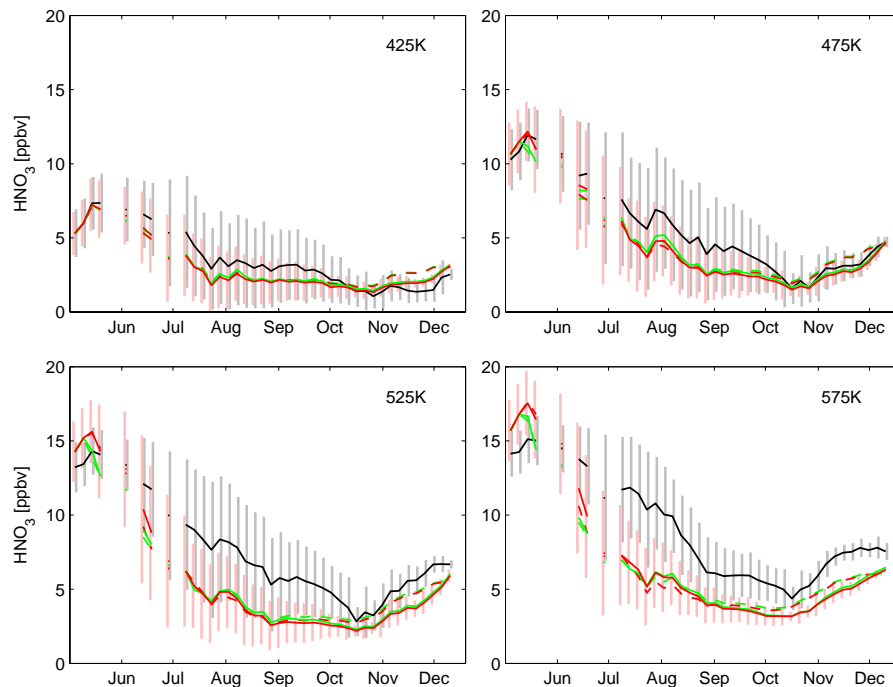


**Fig. 5. (a):** 5-day error-weighted mean vortex 1020 nm extinction of POAM (black) and co-located model extinction contributions: STS (green), NAT (magenta), and ICE (cyan), as well as the total model extinction (red) for the HR100 simulation. Also shown is the model's background aerosol extinction (blue). The SAT extinction does not exceed the axis range. **(b):** zoomed plot of the early winter period showing 5-day error-weighted mean vortex 1020 nm extinction of POAM (black) and the STS (green) and NAT (magenta) contributions to the model extinction for the simulations HR (dashed) and HR100 (full).

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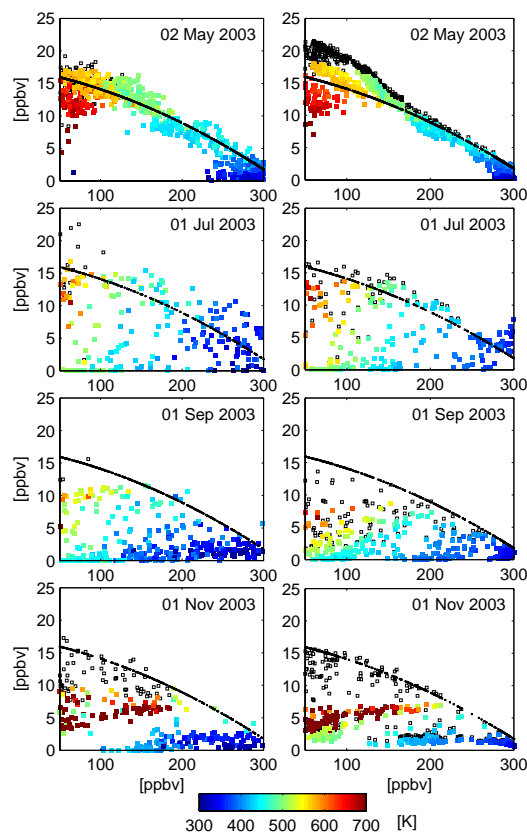


**Fig. 6.** 5-day error-weighted mean vortex  $\text{HNO}_3$  mixing ratios of MIPAS (black) and co-located model results after interpolation of profiles to isentropic levels. Model curves are: LR (green dashed), LR100 (red dashed), HR (green full), HR100 (red full). The grey-shaded area is the 5-day MIPAS variability (1 error-weighted standard deviation) while the red-shaded area is the 5-day variability of the model run HR100.

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**Fig. 7.** Scatterplots of  $\text{HNO}_3$  vs  $\text{N}_2\text{O}$  for MIPAS data inside the polar vortex (left) and the co-located model data (right) for the high resolution version HR100 on four dates. The color-shading is the potential temperature. The small black dots are the relation of Popp et al. (2001). In the MIPAS plots the small black squares are  $\text{HNO}_3 + \text{NO}_2$  and in the model plots the small black squares are the total model  $\text{NO}_y$ .

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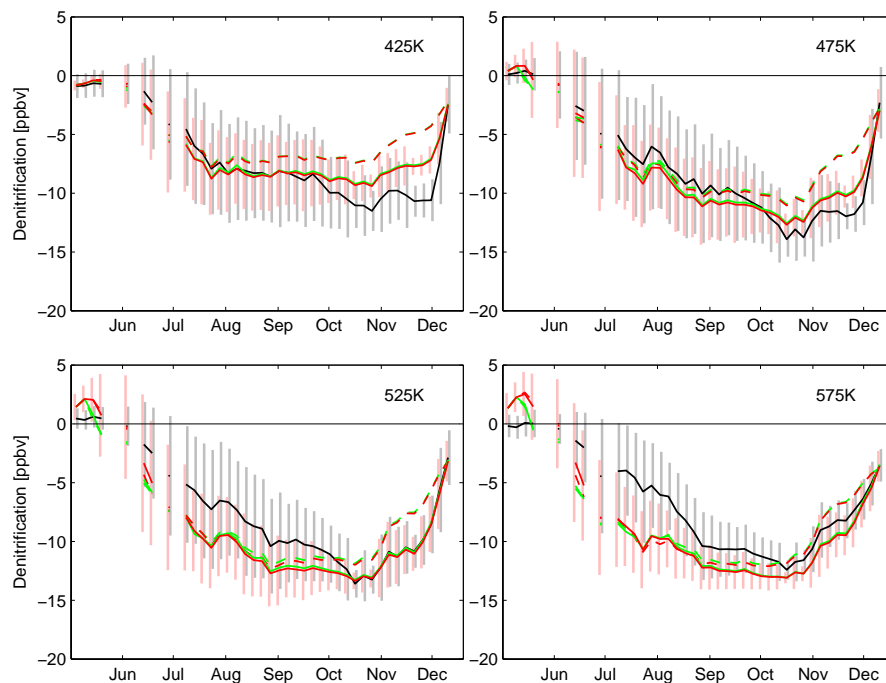
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**Fig. 8.** 5-day error-weighted mean vortex approximative denitrification, calculated as the difference between  $\text{HNO}_3$  and  $\text{NO}_y$  as calculated from  $\text{N}_2\text{O}$  using the relation of Popp et al. (2001), from MIPAS observations (black) and co-located model results after interpolation of profiles to isentropic levels. Color legend as in previous figures.

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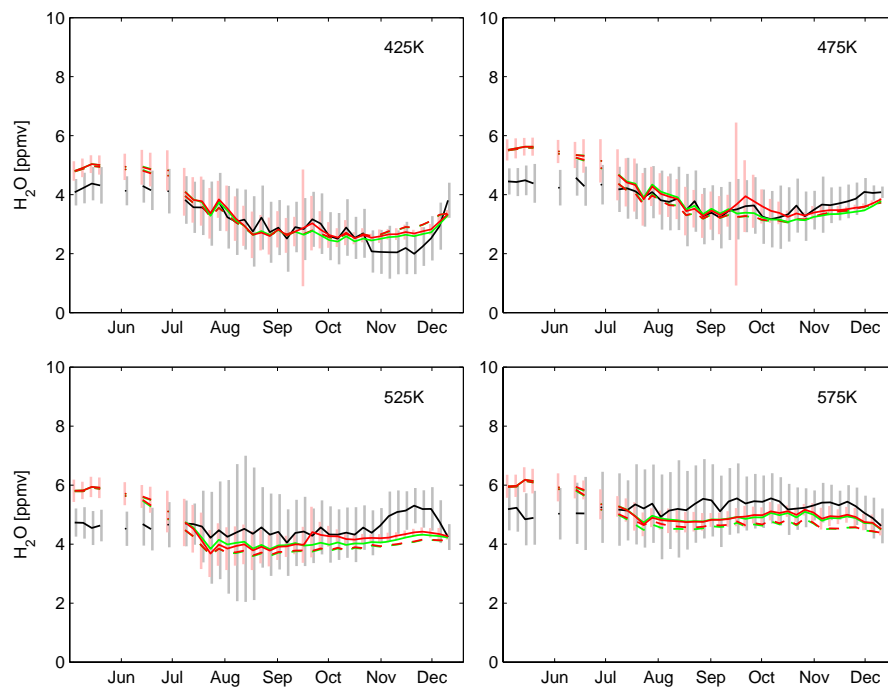
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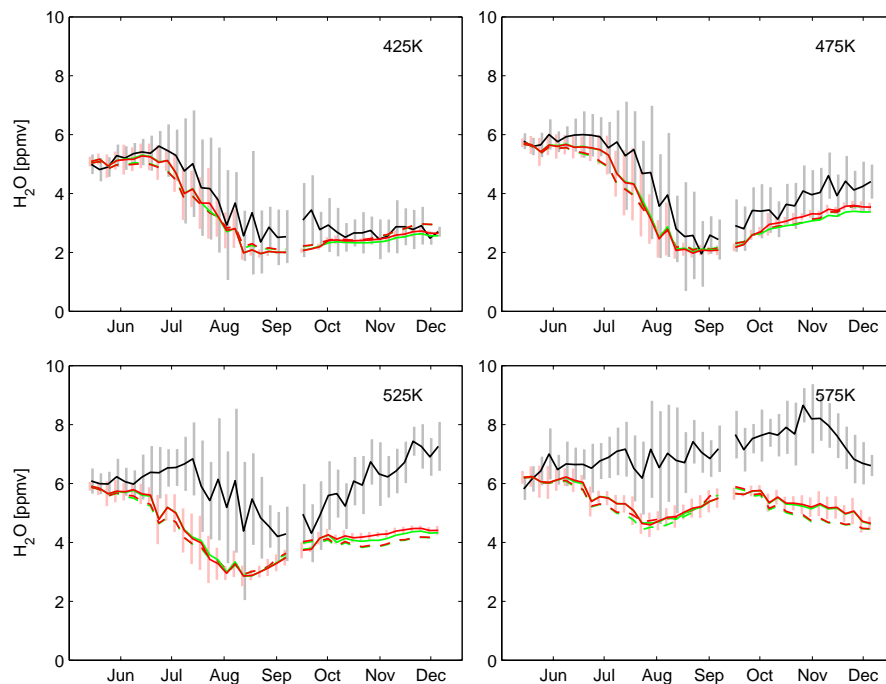
**Fig. 9.** 5-day error-weighted mean vortex  $\text{H}_2\text{O}$  mixing ratios of MIPAS (black) and co-located model results after interpolation of profiles to isentropic levels. Color legend as in previous figures.

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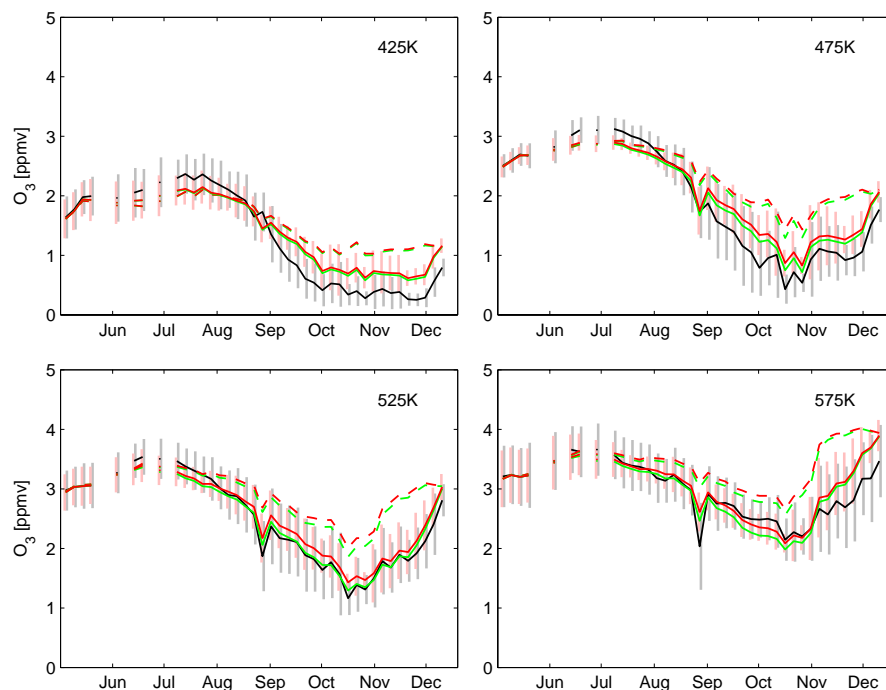


**Fig. 10.** 5-day error-weighted mean vortex  $\text{H}_2\text{O}$  mixing ratios of POAM (black) and co-located model results after interpolation of profiles to isentropic levels. Color legend as in previous figures.

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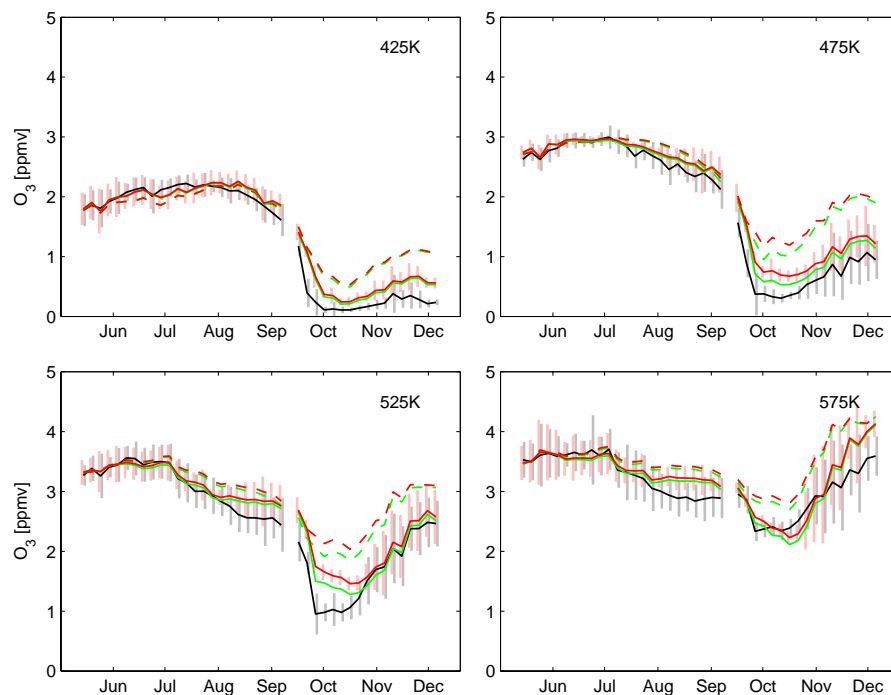


**Fig. 11.** 5-day error-weighted mean vortex  $O_3$  mixing ratios of MIPAS (black) and co-located model results after interpolation of profiles to isentropic levels. Color legend as in previous figures.

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**Fig. 12.** 5-day error-weighted mean vortex  $O_3$  mixing ratios of POAM (black) and co-located model results after interpolation of profiles to isentropic levels. Color legend as in previous figures.

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